

PPS024-03

会場:103

時間:5月22日17:00-17:15

Kaguya 衛星で観測された月周辺の狭帯域ホイッスラー波の特性 Properties of monochromatic whistler waves near the Moon detected by Kaguya

津川 靖基¹*, 寺田 直樹¹, 加藤 雄人¹, 小野 高幸¹, 綱川 秀夫², 高橋 太², 渋谷 秀敏³, 清水 久芳⁴, 松島 政貴² Yasunori Tsugawa¹*, Naoki Terada¹, Yuto Katoh¹, Takayuki Ono¹, Hideo Tsunakawa², Futoshi Takahashi², Hidetoshi Shibuya³, Hisayoshi Shimizu⁴, Masaki Matsushima²

¹ 東北大学大学院理学研究科,² 東京工業大学大学院理工学研究科,³ 熊本大学大学院自然科学研究科,⁴ 東京大学地震研究 所

¹Tohoku University, ²Tokyo Institute of Technology, ³Kumamoto University, ⁴University of Tokyo

We study monochromatic whistler waves near the Moon observed by LMAG aboard the Kaguya spacecraft. The waves were mostly left-hand polarized in the spacecraft frame and observed as narrowband magnetic fluctuations with frequencies close to 1 Hz. Although the monochromatic whistler waves near the Moon were also observed by Lunar Prospector [Halekas et al., 2006], the spectral formation mechanism and the generation process of the waves have not been fully understood.

In the present study, we perform statistical analyses to reveal the properties of the monochromatic whistler waves near the Moon. The results indicate that the intense waves are generated by the solar wind interaction with lunar magnetic anomalies. High occurrence rates of the waves are found in the range of the solar zenith angle (SZA) from 40 ° to 90 ° with remarkable north-south and dawn-dusk asymmetries in the SSE coordinate. Similar asymmetries are found in the selenographical distributions as well. We explain these asymmetries by the localization of magnetic anomalies and the effects of SZA dependencies. Plasma wave theory predicts that the occurrence, intensity, and frequency of the waves are determined by the wave vector direction, magnetic field direction, and solar wind parameters. From the investigation of these parameter spaces we reveal that the properties of the waves can be explained by the Doppler shifted whistler-mode waves propagating against the solar wind.

Moreover, based on the results of the statistical analyses, we compare the properties of the observed waves quantitatively with those of upstream whistler waves from planetary bow shocks reported in previous studies. While we find different properties possibly caused by the differences of the distance from the source region, we discuss the similarities among them suggesting important clues in understanding the generation process of the monochromatic whistler waves near the Moon.

キーワード:ホイッスラーモード,イオン反射,磁気異常

Keywords: whistler-mode, narrowband, magnetic anomaly, reflected ion