

In-situ IR measurement in homogeneous nucleation process of alumina under μG environment

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Homogeneous nucleation process from vapor is characterized by the ratio between time scales for supersaturation increase and for source collision expressed as Λ [1]. Under the physical condition with the same Λ value, homogeneous nucleation process has been regarded to follow the same process. At the dust forming front around evolved stars, Λ value has been calculated to be $\sim 10^{3-5}$ from total pressure and velocity of stellar wind. In contrast, Λ value of $\sim 10^{0-2}$ is known for the gas evaporation method which is one of the simplest experimental methods to produce dust analogues via homogeneous nucleation [2, 3].

In-situ IR measurement during nucleation of nanoparticles in the gas evaporation method proved multi-step formation of metal oxide from vapor to crystalline via liquid droplet in our ground based experiment [4]. Using our advanced technique, we measured IR spectra of nucleating alumina and its evolution while nanoparticles are free-flying under μG environment in which Λ approximates to the value at dust formation region. Specially designed experimental apparatus equipped with dispersive IR spectrometer was loaded to S-520-30 sounding rocket by which the apparatus carried to altitude of 312 km. We also performed ground based experiment combined with FT-IR.

IR spectra of nucleating alumina measured in ground based experiment showed broad absorption extending $>11 \mu\text{m}$. Formed nanoparticles were observed by TEM and identified to δ -alumina. In contrast, sharp absorption centered at $13 \mu\text{m}$ was appeared in μG experiment. This $13 \mu\text{m}$ band is one of the most indicative features of corundum (α -alumina) sphere. Corundum is the most plausible candidate for the origin of unidentified $13 \mu\text{m}$ feature which is often observed for oxygen rich AGB stars with low-mass loss rate [5, 6]. Polymorphic behavior of alumina in homogeneous nucleation process at different Λ will be the key to understand astronomical dust formation.

Reference

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