

Detection of crystalline silicate in T Tauri-type stars by mid-infrared spectroscopic observation

Mitsuhiro Honda[1], Takuya Yamashita[2], Hirokazu Kataza[3], Takashi Miyata[4], Yoshiko Okamoto[5], Shigeyuki Sako[1], Shinya Takubo[1], Takashi Onaka[1]

[1] Department of Astronomy, University of Tokyo, [2] National Astronomical Observatory, Japan, [3] ISAS, [4] Institute of Astronomy, University of Tokyo, [5] Institute of Physics, Kitasato University

Crystalline silicates have already been observed in comets, IDPs, beta Pic (Vega-like star), though ISM silicates and younger objects such as protostars, T Tauri stars and Herbig Ae/Be stars have not. In recent years, ISO and other ground-based instruments showed that some Herbig Ae/Be stars possess crystalline silicates, which imply some crystallization event must occur at Herbig Ae/Be stage in intermediate stars (2 to 10 solar-mass stars). But low-mass young stellar objects such as T Tauri stars show no evidence of existence of crystalline silicates, so when crystallization of silicates has occurred in low-mass stars that is believed to be a precursor of our Solar-System, is an interesting matter.

We have carried out mid-infrared spectroscopic observations of older T Tauri stars using COMICS with SUBARU, and for the first time, we have detected crystalline silicates in some T Tauri stars.

We will discuss it further from the view of planet formation and try to compare with our early Solar-System formation events.