

On the bi-polar magnetic structure at the leading-edge of reconnection jets

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A recent paper by Slavin et al. [2003] shows that some of the jets in the magnetotail have a bi-polar magnetic structure at their leading edges. It is characterized by southward-then-northward variation for earthward jets, and vice versa.

In this paper, we try to model this feature in terms of three-dimensional reconnection with non-zero guide field.

Here the coordinates system is the conventional one used for magnetotail studies.

As the initial condition for the three-dimensional MHD simulations, we set a Harris current sheet $B_x = \tanh(z)$ with uniform guide field $B_y = B_{y0}$. Reconnection is initiated by putting an ad-hoc anomalous resistivity in a localized region that has a finite extent in the y (dawn-dusk) direction as well, whose half-width is denoted by R_y . Such a three-dimensional reconnection in the presence of the guide field results in the leading-edge structure as follows:

- (1) The jet is bifurcated to higher-latitudes at the leading edge.
- (2) When B_y is non-zero, unlike the $B_y=0$ cases, the bifurcated jet is rotated in the yz plane. When B_y is positive and for the earthward propagating part, hot and fast flowing plasma is located at north-dusk and south-dawn quadrants.
- (3) Upon arrival of the bifurcated jet leading part, plasma flow in the yz plane is generated as the two spots of hot plasma try to expand.
- (4) The flow in the yz plane bends the magnetic field line at the equator such that the B_z component is negative at the midnight meridian.
- (5) In the earthward jet itself, the B_z component is positive.

Data from a spacecraft situated at the equator over which this earthward jet passes would record southward-then-northward variation in the magnetic field, which is indeed reported by Slavin et al.

The behavior of the B_y component, the V_x component, and the plasma density also show good qualitative agreement with the data.

These hold the same when B_y is negative.

To make an quantitative assessment, we have focused on the minimum B_z during the events, which is reported to be -3 nT on average, and explored in what (B_{y0}, R_y) -space the initial condition has to be to have the minimum B_z to be equal to this value.

For $B_{y0}=4$ nT, which is the reported typical guide field strength, we find that R_y must be 4 times the half current sheet thickness. Taking the half-thickness to be 1,000-3,000 km, the dawn-dusk width of the reconnection jet ($2R_y$) is determined to be 1-4 R_e .