

# Surface reactions in interstellar clouds: Rate constants of hydrogenation reactions and diffusion constants of hydrogen atoms

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The observations of interstellar ices or comets detected formaldehyde (H<sub>2</sub>CO) and methanol (CH<sub>3</sub>OH) (Crovisier & Bockelee-Morvan 1999, *Space Sci Rev.* 90, 19; Keane et al. 2001, *A&A* 376, 254). Several kinds of reactions to form H<sub>2</sub>CO and CH<sub>3</sub>OH have been proposed such as gas-phase reactions (Shalabiea & Greenberg 1994, *A&A* 290, 266), UV photolysis on H<sub>2</sub>O+CO ices (Schutte et al. 1996, *A&A* 309, 633), proton bombardment on H<sub>2</sub>O+CO ices (Hudson & Moore 1999, *Icarus* 140, 451), and hydrogenation on H<sub>2</sub>O+CO ices (Tielens & Whittet 1997, in *Molecules in Astrophysics: Probe and Processes*, 45).

Hydrogenation is proposed to be the one of the most effective reactions (Tielens & Whittet 1997, in *Molecules in Astrophysics: Probe and Processes*, 45), and laboratory experiment has been carried out recently (Hiraoka et al. 2002, *ApJ* 577, 265; Watanabe & Kouchi 2002, *ApJ* 571, L173).

We analyze the data of the quantitative hydrogenation experiment of CO by Watanabe & Kouchi (2002, *ApJ* 571, L173), Watanabe et al. (2002, *ApJ* 588, L121), and Hidaka et al. (2003, *Proc. 36th ISAS LPS*, 250), and derive the rate constants of hydrogenation of CO and H<sub>2</sub>CO and the diffusion constant of H atoms in the ice.