

# Oxygen isotopic compositions of refractory inclusions in CH chondrite

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Oxygen isotope systematics between  $^{16}\text{O}$ -rich and  $^{16}\text{O}$ -poor reservoirs is well known in the solar system (Clayton, 1993). Oxygen isotopic composition of component of chondrite of  $\delta^{17}\text{O}$ ,  $\delta^{18}\text{O}$  relative to SMOW is plotted on CCAM line or slope-1 line (Young and Russell, 1998).

Refractory inclusions (Ca, Al-rich inclusions (CAIs) and amoeboid olivine aggregates (AOAs)) in a polished thin section of Acfer214 CH chondrite were studied by BSE imaging, X-ray elemental mapping and EPMA equipped with Oxford LINK ISIS EDS. Oxygen isotopes were measured in-situ by ion microprobe spot analysis ( $^{16}\text{O}$ ,  $^{17}\text{O}$  and  $^{18}\text{O}$ ) and ion microscope ( $^{16}\text{O}$  and  $^{18}\text{O}$ ) by TiTech Cameca ims1270 SIMS instrument equipped with SCAPS (2-D ion detector).

CAIs in Acfer214 are mineralogically classified into (1) grossite-melilite [Al-diopside rim] inclusions, (2) grossite (hibonite)-spinel-melilite [Al-diopside + olivine rim] inclusions, (3) spinel-melilite [Al-diopside rim] inclusions, (4) pyroxene spherule and (5) pyroxene-spinel [Al-diopside rim] spherule and otherwise  $\text{CaAl}_2\text{O}_4$  were found in only one inclusion and the inclusion seems to be corroded. AOAs are composed of olivine, pyroxene, spinel and anorthite.

Oxygen isotopic compositions of (1), (2), (3) and (5) were measured by ion microprobe spot analysis, which are distributed continuously on slope-1 line ( $\delta^{17}\text{O}$ ,  $\delta^{18}\text{O}$  SMOW = -60 - 0 permil) on a 3 isotope diagram. (1) and (3) are relatively  $^{16}\text{O}$ -rich and (2) and (5) are relatively  $^{16}\text{O}$ -poor. Results of  $\delta^{18}\text{O}$  mapping of some CAIs and AOAs show that they have nearly homogeneous oxygen isotopic compositions. (4) and AOAs are  $^{16}\text{O}$ -rich ( $\delta^{18}\text{O}$  = -30 permil and -50 - -30 permil, respectively) and  $\text{CaAl}_2\text{O}_4$  is  $^{16}\text{O}$ -poor (nearly terrestrial value).

Matrix of Acfer214 seems to be altered by terrestrial weathering, but refractory inclusions are primitive because they have no alteration products and no oxygen isotopic alteration. Oxygen isotopic compositions of refractory inclusions reflect formation condition, place and time. Wide distribution of oxygen isotopic compositions shows that refractory inclusions in Acfer214 were formed in different places, time, conditions and accreted into Acfer214 parent body.