Precipitating electron beams observed below auroral acceleration region

Kazushi Asamura[1]; Takeshi Sakanoi[2]; Yusuke Ebihara[3]; Kanako Seki[4]; Masafumi Hirahara[5]; Yasuyuki Obuchi[6]; Tomohiro Ino[2]; Yasumasa Kasaba[7]

[1] ISAS/JAXA; [2] PPARC, Grad. School of Sci., Tohoku Univ.; [3] Nagoua Univ., IAR; [4] STEL, Nagoya Univ.; [5] Department of Physics, Rikkyo University; [6] Planet Plasma Atmos, Tohoku Univ; [7] JAXA/ISAS

Reimei satellite is polar orbiting satellite within altitude range of 610 to 670km, which is lower than auroral acceleration regions but higher than auroral emission regions. The satellite is three-axes stabilized, we cannot apply typical techniques of getting three-dimensional particle distribution functions with satellite spin motion. However, we can get particle energy spectra with full pitch-angle coverage, and with time reslution of 40ms, since auroral particle instruments (ESA/ISA) onboard are top-hat type analyzers, and installed into appropriate positions.

Precipitating electrons with energies of 10 to 1000eV are frequently observed in high-latitude side of auroral 'Inverted-V' structures with active aurora. These electrons are highly confined field-aligned beams which are time-dispersed. They would be cold ionospheric electrons accelerated by inertial Alfven waves which are generated in higher altitude and traveling downward.

The ionospheric cold electrons can get only ~300eV due to the inertial Alfven waves. Resulting auroral photon emission would be faint. Therefore, auroral emission structures would almost be determined by accelerated electrons through the inverted-Vs. We picked up events of the time-dispersed precipitating electrons without the inverted-V signatures from Reimei particle data. From auroral image data taken by MAC onboard Reimei, corresponding auroral structures are line-like (thickness of a few km) with traveling voltex.