

The role of ionospheric O^+ ions in triggering substorms

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In the literature the role of ionospheric O^+ ions in triggering substorms has been discussed. Some of them insisted that plasma sheet instabilities, which trigger a substorm, are likely to arise when the O^+ ion density is high.

The purpose of this study is to investigate the role of O^+ ions in the plasma sheet by examining whether or not a substorm occurs in the region where O^+ ions are dominant components. In this research we used energetic ion flux data collected by the energetic particles and ion composition (EPIC) instrument on board the GEOTAIL spacecraft. We analyze 5-year data (from 2000 to 2004) in the plasma sheet ($X_{GSM} = -8$ to $-23R_E$, $Y_{GSM} = -15$ to $15 R_E$). Substorm events are selected from the IMAGE/FUV substorm onset list (Frey et al., JGR, 2004).

We examine the MLT distribution of substorm onsets (auroral brightenings) when the ion number density ratio (O^+/H^+) is high at a given region. Then, we compare the distribution with that when the ratio is low at the same region. From this analysis we obtain the following results:

1) Substorm onsets, viz. auroral brightenings, are likely to occur in the dusk (dawn) side when the ion number density ratio of O^+/H^+ is high in the dusk (dawn) side.

2) The abovementioned property is predominantly observed in the near-Earth plasma sheet (especially where $X = -8$ to $-11R_E$).