

On the plasma sheet variations at the time of substorm onset: A case study of THEMIS observations

Shinobu Machida[1]; Yukinaga Miyashita[2]; Akimasa Ieda[2]; Masahito Nose[3]; Vassilis Angelopoulos[4]; James P. McFadden[4]; H. Uli Auster[5]

[1] Dept. of Geophys., Kyoto Univ.; [2] STEL, Nagoya Univ.; [3] DACGSM, Kyoto Univ.; [4] SSL, UC Berkeley; [5] TUBS

<http://www-step.kugi.kyoto-u.ac.jp/~machida/>

In a series of our studies, the time development of Earth's magnetosphere during substorms has been investigated by adopting a superposed-epoch analysis to the Geotail data. In those studies, time variations during substorms were obtained for the earthward flow velocity, northward magnetic field, total pressure, electric field, and Poynting flux toward the center of the plasma sheet etc. We could confirm various variations which relevant models of substorm are based on or predict. However, none of them could perfectly explain our results. Therefore we proposed a new model called Catapult (Slingshot) Current Sheet Relaxation model. In this model, the initial variation starts at $X \sim 14$ Re in the magnetotail rather than $X \sim 8$ Re as predicted by the current disruption model, or $X \sim 20$ Re as predicted by the near-Earth neutral line model.

During the growth phase of substorm, the Poynting flux toward the plasma sheet center enhances the cross-tail current. Then the highly stretched current sheet with small dipole field relaxes resulting in the earthward motion of that current sheet. Simultaneously, the reduction of the total pressure progresses around the flow region at $X \sim 16$ Re, and the Poynting flux toward the plasma sheet center further enhances about 4 min before the auroral breakup. Those variations work as a seed for the disturbances in the inner magnetosphere which was in a marginal state against the instability such as ballooning instability. The occurrence of the instability produces the current disruption and dipolarization of the magnetic field. Meanwhile, the earthward flows produce a very thin current sheet at the boundary between the current sheet with highly stretched dipole field lines and the Harris-type current sheet at $X \sim 20$ Re. The magnetic reconnection starts in this very thin current sheet located at $X \sim 20$ Re.

Based on those results, we performed a case study of substorm events observed by THEMIS spacecraft. When the onset was observed at ~ 0811 UT on March 21, 2008, five THEMIS spacecraft were located in the region from -8 Re to -15 Re in the X(GSM)-direction, and from 5 Re to 7 Re in the Y(GSM)-direction. A sudden increase of the northward magnetic field (B_z) was observed at 0808UT by P1 spacecraft located at $X \sim 15$ Re and P2 located at $X \sim 12$ Re. About 1 min later, a similar increase of B_z was observed by neighboring P3 and P4 spacecraft located at $X \sim 10$ Re. A gradual increase of B_z was observed at 0811UT at $X \sim 8$ Re by P5 spacecraft located at $X \sim 8$ Re. All events were accompanied by the earthward flows and the reduction of the total pressure (the sum of the plasma pressure and magnetic pressure). The variations detected by each spacecraft were almost synchronized with the increase of B_z . These characters were also found in common in the other substorm subsequently occurred at 0928UT.

Those characteristic variations propagated from the region of $X \sim 15$ Re to $X \sim 8$ Re, which apparently contradicts with that predicted by the current disruption model in which the current disruption and associated total pressure reduction first occurs at $X \sim 8$ Re and propagates as a form of rarefaction wave toward the down tail region. The variations in the magnetic field, the flow and the total pressure start a few minutes before the substorm onset in the region of $X \sim 14$ Re, which strongly supports the validity of the catapult (slingshot) current sheet relaxation model.