

## The origin of N<sub>2</sub> in Titan's atmosphere: the role of impact devolatilization of Titan's icy crust

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Titan has a thick atmosphere composed primarily of N<sub>2</sub>. One of the most puzzling aspects of Titan's atmosphere is the origin of N<sub>2</sub> because of the near absence of non-radiogenic noble gases in the atmosphere, highly suggestive of that the nitrogen was captured as NH<sub>3</sub> and other non-N<sub>2</sub>-bearing compounds in the satellitesimals. Although several studies have investigated the mechanism responsible for converting NH<sub>3</sub> to N<sub>2</sub> in the primitive atmosphere of Titan [Mckay et al 1988, Atreya et al 1978], it is still unclear how and when the production of N<sub>2</sub> has occurred.

In this study, we assess the role of shock-induced devolatilization of Titan's icy crust by hypervelocity impacts of cometary bodies for the origin of N<sub>2</sub> in the atmosphere. Although the chemical compositions of Titan's crust are still uncertain, NH<sub>3</sub>-ice [Nelson et al 2008] and ammonium sulfate [Fortes et al 2007] are considered as major components. To investigate whether the conversion of NH<sub>3</sub>-ice and ammonium sulfate to N<sub>2</sub> occur or not by cometary impacts, we conducted laboratory experiments of hypervelocity impact using a laser gun. We used gold (Au) foils, varied thickness to vary impact velocity, as impactors, and isotopic-labeled(<sup>15</sup>N) ammonium sulfate and NH<sub>3</sub>-H<sub>2</sub>O ice (50% of NH<sub>3</sub>) as targets.

We experimentally obtain the efficiency of <sup>15</sup>N<sub>2</sub> production as a function of the peak shock pressure. We found that N<sub>2</sub> production begins at about 10 GPa of the peak shock pressure, and that the efficiency of N<sub>2</sub> production linearly increases with the peak shock pressure. When considering the peak shock pressure achieved in Titan's crust by impacts of planetesimals and comets, our experimental results suggest that N<sub>2</sub> production from ammonium sulfate and NH<sub>3</sub>-ice in Titan's crust takes place efficiently in these impacts. Furthermore, using our experimental data of the efficiency of N<sub>2</sub> production, we estimate the total amount of N<sub>2</sub> produced by cometary impacts over 4.5 Gyr on Titan. The estimated value of N<sub>2</sub> production in our study shows that a large part of or almost all the present amount of N<sub>2</sub> could have been derived from the devolatilization of Titan's crust by cometary impacts.