

PPS003-P01

会場:コンベンションホール

時間:5月26日 10:30-13:00

## 月隕石 NWA 4734 に記録された衝突履歴の研究 Dynamic event recorded in a lunar meteorite NWA 4734

金子 詳平<sup>1\*</sup>, 大谷 栄治<sup>1</sup>, 宮原 正明<sup>1</sup>, 境 毅<sup>2</sup>, 鹿山 雅裕<sup>3</sup>, 西戸 裕嗣<sup>3</sup>, 大石 泰生<sup>4</sup>, 平尾 直久<sup>4</sup>

Shohei Kaneko<sup>1\*</sup>, Eiji Ohtani<sup>1</sup>, Masaaki Miyahara<sup>1</sup>, Takeshi Sakai<sup>2</sup>, Masahiro Kayama<sup>3</sup>, Hirotsugu Nishido<sup>3</sup>, Yasuo Oishi<sup>4</sup>, Naohisa Hirao<sup>4</sup>

<sup>1</sup> 東北大学大学院理学研究科地学専攻, <sup>2</sup> 東北大学国際高等研究教育機構国際高等融合研, <sup>3</sup> 岡山理科大学自然科学研究所, <sup>4</sup> 財団法人高輝度光科学研究センター

<sup>1</sup>Tohoku University, <sup>2</sup>IIAIR, Tohoku University, <sup>3</sup>Okayama University of Science, <sup>4</sup>JASRI

It is assumed that Late Heavy Bombardment (LHB) occurred in the inner solar system from 4.1 to 3.8 billion years ago. Many planetesimal and/or meteorites collided on the Earth during LHB. The Moon also would experience LHB because many impact craters and thick regolith exist on the Moon. Lunar meteorites would have information about such dynamic events occurred on the Moon. Shock pressure condition could be estimated based on high-pressure mineral assemblages contained in the lunar meteorites.

In this study, we investigated a lunar meteorite, Northwest Africa (NWA) 4734 by scanning electron microscope (SEM), Raman spectroscopy, electron probe micro analyzer (EPMA), X-ray diffraction (XRD) and Cathodoluminescence (CL) spectroscopy. Previous work [1] reports that the high-pressure polymorph of SiO<sub>2</sub> may exist in NWA 4734. Accordingly, we focused our investigation on the high-pressure polymorphs of SiO<sub>2</sub>.

NWA 4734 is unbrecciated basalt. Major constituent minerals of NWA 4734 studied here are clinopyroxene, plagioclase (maskelynite), olivine, ilmenite, and SiO<sub>2</sub>. Many shock-melt veins and melt pockets exist in the NWA 4734. Raman spectroscopy and SEM observations show that dendritic coesite exists in the SiO<sub>2</sub> grain entrained in the shock-melt vein. Raman spectroscopy indicates that pyroxene-glass was identified in the matrix of the shock-melt vein, which might originate from silicate-perovskite. SiO<sub>2</sub> grains with lamellar textures were observed in a host-rock of NWA 4734. XRD patterns indicate that the SiO<sub>2</sub> grain contains alpha-PbO<sub>2</sub>-type SiO<sub>2</sub> seifertite (a = 4.079(2) Å, b = 5.030(2) Å, c = 4.485(1) Å). A small amount of stishovite was identified by XRD, CL and Raman in some SiO<sub>2</sub> grains with lamellar textures. The phase equilibrium diagram of SiO<sub>2</sub> show that cristobalite transforms to seifertite at ~40 GPa at room temperature [2, 3]. It is likely that the SiO<sub>2</sub> grain was originally cristobalite based on several previous studies [e.g., 1, 4]. Accordingly, shock pressure condition recorded in NWA 4734 is at least ~40 GPa.

### Reference

[1] H. Chennaoui Aoudjehane, A. Jambon, First evidence of high pressure silica: stishovite and seifertite in lunar meteorite Northwest Africa 4734, 71st Annual Meteoritical Society Meeting (2008) (abstract #5058).

[2] Y. Tsuchida, T. Yagi, New pressure-induced transformations of silica at room temperature, *Nature* 347 (1990) 267-269.

[3] L. S. Dubrovinsky, N. A. Dubrovinskaia, S. K. Saxena, F. Tutti, S. Rekhi, T. Le Bihan, Guoyin Shen, J. Hu, Pressure-induced transformations of cristobalite, *Chemical Physics Letters* 333 (2001) 264-270.

[4] J. J. Papike, G. Ryder, C. K. Shearer, Lunar samples, In *Planetary Materials, Reviews in Mineralogy & Geochemistry* 36, edited by J. J. Papike (1998) 5-1-5-234.

キーワード: 月, シリカ, コーサイト, スティショバイト, ザイフェルタイト

Keywords: moon, silica, coesite, stishovite, seifertite