

U005-05

会場:国際会議室

時間:5月26日09:30-09:50

はやぶさ回収試料の初期分析:イトカワレゴリス粒子の3次元構造と初期分析にお けるX線マイクロトモグラフィーの役割 Preliminary examination of Hayabusa asteroidal samples: 3-D structures of Itakawa par-

ticles using X-ray microtomography

土山 明^{1*}, 上椙 真之², 上杉 健太朗³, 中野 司⁴, 中村 智樹⁵, 野口 高明⁶, 野口 遼¹, 松本 徹¹, 松野 淳也¹, 永野 宗¹, 竹内 晃久³, 鈴木 芳生³, 海老原 充⁷, 今井 悠太¹, トレバー アイルランド⁸, 北島 富美雄⁹, 松島 亘志¹⁰, 道上 達広¹¹, 長尾 敬 介¹², 奈良岡 浩⁹, 岡崎 隆司⁹, スコット サンドフォード¹³, 圦本 尚義¹⁴, マイケル ゾレンスキー¹⁵, 藤村 彰夫¹⁶, 安部 正 真¹⁶, 矢田 達¹⁶, 向井 利典¹⁶, 岡田 達明¹⁶, 石橋 之宏¹⁶, 白井 慶¹⁶, 上野 宗孝¹⁶, 川口 淳一郎¹⁶, 吉川 真¹⁶

Akira Tsuchiyama^{1*}, Masayuki Uesugi², Uesugi Kentaro³, Tsukasa Nakano⁴, Tomoki Nakamura⁵, Takaaki Noguchi⁶, Ryo Noguchi¹, Tooru Matsumoto¹, Junya Matsuno¹, Takashi Nagano¹, Takeuchi Akihisa³, Suzuki Yoshio³, Mitsuru Ebihara⁷, Yuta Imai¹, Trevor R. Ireland⁸, Fumio Kitajima⁹, Matsushima Takashi¹⁰, Tatsuhiro Michikami¹¹, Keisuke Nagao¹², Hiroshi Naraoka⁹, Ryuji Okazaki⁹, Scott A. Sndford¹³, Hisayoshi Yurimoto¹⁴, Michael E. Zolensky¹⁵, Akio Fujimura¹⁶, Masanao Abe¹⁶, Toru Yada¹⁶, Toshifumi Mukai¹⁶, Tatsuaki Okada¹⁶, Yukihiro Ishibashi¹⁶, Kei Shirai¹⁶, Munetaka Ueno¹⁶, Junichiro Kawaguchi¹⁶, Makoto Yoshikawa¹⁶

¹ 大阪大学理学研究科,² 大阪大学工学研究科,³ 高輝度光科学研究センター SPring8,⁴ 産業技術総合研究所,⁵ 東北大学理 学研究科,⁶ 茨城大学理学部,⁷ 首都大学東京理工学研究科,⁸ オーストラリア国立大学,⁹ 九州大学理学府,¹⁰ 筑波大学シス テム情報工学研究科,¹¹ 福島工業高等専門学校,¹² 東京大学理学系研究科,¹³NASA Ames 研究センター,¹⁴ 北海道大学理 学研究院,¹⁵NASA Johnson 研究センター,¹⁶ 宇宙航空研究開発機構

¹Graduate School of Sci., Osaka Univ., ²Graduate School of Eng., Osaka Univ., ³JASRI, SPring-8, ⁴AIST, ⁵Graduate School of Sci., Tohoku Univ., ⁶Coll. of Sci. at Ibaraki Univ., ⁷Grad. Sch. Sci. Eng., Tokyo Met. Univ., ⁸Res. School of Earth Sci., ANU, ⁹Graduate School of Sci., Kyushu Univ., ¹⁰Grad. Sch. Sys. Inf. Eng., Tsukuba Univ., ¹¹Fukushima National College of Technology, ¹²Graduate School of Sci., Univ. of Tokyo, ¹³NASA Ames Research Center, ¹⁴Graduate School of Sci., Hokkaido Univ., ¹⁵NASA Johnson Research Center, ¹⁶JAXA

Particles on S-type Asteroid 25143 Itokawa were successfully recovered by the Hayabusa mission of JAXA [1]. This is the first regolith sample of asteroid. Forty-one particles of 30-150 microns has been allocated to mainstream flow of preliminary examination (PE) [2]. These particles were imaged using SR -based X-ray mirotomography to obtain their three-dimensional structures as the first analysis in sequential mainstream analyses.

X-ray absorption imaging tomography system was used at BL47XU of SPring-8, Hyogo, Japan [3]. Each particle was imaged using two X-ray energies at 7 and 8 keV. As these energies are smaller and larger than the K?adsorption edge of Fe (7.11 keV), respectively, we can recognize olivine, Ca-poor pyroxene, Ca-rich pyroxene, plagioclase, troilite, taenite, kamacite, chromite and Ca phosphate easily using the contrasts of CT images, which correspond to linear attenuation coefficient (LAC) obtained by tomographic reconstruction. Relatively smaller thirty-five particles were imaged with the voxel (pixel in 3-D) size of about 90 nm (effective spatial resolution: 300-400 nm), while 6 particles with the voxel size of about 210-250 nm (effective spatial resolution of > 400 nm).

3-D internal structures (3-D distribution of LAC values) of the particle were obtained. Based on the LAC values, we recognized most of the mineral phases and their three-dimensional distribution. Nineteen particles are mono-mineralic or almost mono-mineralic, fifteen particles are poly-mineralic (some of them contains accessory minerals), and seven particles have specific textures, such as a particle containing a porous aggregate of sub-micron grains. Voids are usually observed. As the particle size is comparable to the typical mineral grain sizes of LL5 and 6 chondrites, candidates of Itokawa surface sample, information of 3-D volume will help to reconstruct original textures of Itokawa materials.

From now on, modes of the minerals will be obtained. By using and image analysis, the external shapes of the particles will be also extracted to obtain volume, porosity and three-axial lengths. These shape parameters will be compared with those of boulders on Itokawa [4], fragments of collisional experiments [5] and lunar regolith [6] to obtain information on formation process of Itokawa regolith. By combining the chemical compositions of the minerals [7], the chemical composition and density of each particle and bulk composition and density will be also obtained.

After the tomographic imaging, the samples moved on to the non-destructive XRD and XRF analysis at KEK-PF and SPring-8, and then will move on to destructive analyses, such as those by SEM, EPMA, SIMS and TEM. Combination of the non-destructive analyses, microtomography, XRD and XRF, used for design of the later destructive analyses is one of the key features in this

PE. 3-D phase map gives information where should be cut for the later destructive analyses to obtain suitable areas of specific minerals in cross sections of small particles. Carbonaceous materials in a particle can be identified by the microtomography if they are present.

[1] Nakamura T. (2011) abstract in 42nd LPSC. [2] Tsuchiyama A. et al. (2011) abstract in this conference. [3] Uesugi K. et al. (2006) Proc. SPIE, 6318, 63181F. [4] Michikami T. et al. (2010) Icarus 207, 277, [5] Fujiwara et al., (1978) Nature, 272, 602.
[6] Matsushima T. et al. (2009) J. Aerospace Eng., 22, 1, 15-23. [7] Nakamura T. (2011) abstract in this conference.

キーワード: はやぶさ計画, 小惑星, レゴリス, スプリング8, 3次元構造, X線 CT Keywords: Hayabusa mission, asteroid, regolith, SPring-8, three-dimensional structure, X-ray tomography