

The Magnetic and Shielding Effects of Ring Current on Radiation Belt Dynamics The Magnetic and Shielding Effects of Ring Current on Radiation Belt Dynamics

Mei-Ching Fok^{1*}

FOK, Mei-Ching^{1*}

¹NASA Goddard Space Flight Center

¹NASA Goddard Space Flight Center

The ring current plays many key roles in controlling magnetospheric dynamics. A well-known example is the magnetic depression produced by the ring current, which alters the drift paths of radiation belt electrons and may cause significant electron flux dropout. Little attention is paid to the ring current shielding effect on radiation belt dynamics. A recent simulation study that combines the Comprehensive Ring Current Model (CRCM) with the Radiation Belt Environment (RBE) model has revealed that the ring current-associated shielding field directly and/or indirectly weakens the relativistic electron flux increase during magnetic storms. In this talk, we will discuss how ring current magnetic field and electric shielding moderate the radiation belt enhancement.

キーワード: Radiation Belts, Ring Current, magnetic Storm, Relativistic electrons

Keywords: Radiation Belts, Ring Current, magnetic Storm, Relativistic electrons

Simultaneous Global Observations of Ring Current Dynamics Simultaneous Global Observations of Ring Current Dynamics

Eric Grimes^{1*}, Joseph D. Perez¹, Jerry Goldstein², David J. McComas², Phil Valek²
GRIMES, Eric^{1*}, Joseph D. Perez¹, Jerry Goldstein², David J. McComas², Phil Valek²

¹Department of Physics, Auburn University, AL, USA, ²Southwest Research Institute, San Antonio, TX, USA

¹Department of Physics, Auburn University, AL, USA, ²Southwest Research Institute, San Antonio, TX, USA

Studying the structure and evolution of magnetospheric plasmas is fundamental to understanding our near Earth space environment. The Two Wide-angle Imaging Neutral-atom Spectrometers (TWINS) mission provides a unique opportunity to image global ring current plasmas by measuring energetic neutral atoms (ENAs) created when ions trapped on geomagnetic field lines undergo charge exchange with cold neutral hydrogen. The ion distribution can be deconvolved from the ENA images using a modern inversion technique. In this talk, we use data from a variety of orbiting spacecraft to investigate ring current ion dynamics during the main and early recovery phases of several recent ICME-driven geomagnetic storms. Global proton distributions deconvolved from simultaneous TWINS ENA observations are used to provide context to GOES and THEMIS in-situ energetic particle and magnetometer observations of the ring current. Several aspects of inner magnetospheric dynamics are quantified from a global perspective, including equatorial ion pitch angle anisotropy, density and temperature throughout a typical ICME-driven storm.

キーワード: ring current dynamics, proton pitch angle anisotropy, inner magnetosphere, energetic neutral atoms
Keywords: ring current dynamics, proton pitch angle anisotropy, inner magnetosphere, energetic neutral atoms

Pc1/EMIC waves observed at subauroral latitude during sudden magnetospheric compressions

Pc1/EMIC waves observed at subauroral latitude during sudden magnetospheric compressions

Khan-Hyuk Kim^{1*}, Kazuo Shiokawa², Dong-Hun Lee¹, Hyuck-Jin Kwon¹, Martin Connors³
KIM, Khan-Hyuk^{1*}, Kazuo Shiokawa², Dong-Hun Lee¹, Hyuck-Jin Kwon¹, Martin Connors³

¹School of Space Research, Kyung Hee University, South Korea., ²Solar-Terrestrial Environment Laboratory, Nagoya University, Nagoya, Japan., ³Centre for Science, Athabasca University, Canada.

¹School of Space Research, Kyung Hee University, South Korea., ²Solar-Terrestrial Environment Laboratory, Nagoya University, Nagoya, Japan., ³Centre for Science, Athabasca University, Canada.

It is generally accepted that sudden compressions of the magnetosphere cause electromagnetic ion cyclotron (EMIC) wave growth by increasing the proton temperature anisotropy. These compression-associated EMIC waves are expected to be on higher latitudes (i.e., higher-L regions close to the magnetopause). In this study we examine Pc1 pulsations, which are believed to be generated by the EMIC instability, observed at subauroral latitude near the nominal plasmapause when the magnetosphere is suddenly compressed by solar wind dynamic pressure variations, using induction magnetometer data obtained from Athabasca, Canada (geomagnetic latitude = 61.7deg N, L ~ 4.5). We identified 9 compression-associated Pc1 waves with frequencies of ~0.5-2.0 Hz. The wave activity appears in the horizontal H (positive north) and D (positive eastward) components. All of events show low coherence between H and D components. This indicates that the Pc1 pulsations in H and D oscillate with a different frequency. Thus, we cannot determine the polarization state of the waves. We will discuss the occurrence location of compression-associated Pc1 pulsations, their spectral structure, and wave properties.

キーワード: EMIC, Pc1, Magnetospheric compression, Temperature anisotropy, Plasmapause, Subauroral latitude
Keywords: EMIC, Pc1, Magnetospheric compression, Temperature anisotropy, Plasmapause, Subauroral latitude

GEMSIS-RC モデルに基づいた環電流とプラズマ圏が内部磁気圏 ULF 波動に与える影響の研究

Effects of the ring current and plasmasphere on ULF waves in the inner magnetosphere based on the GEMSIS-RC model

関 華奈子^{1*}, 天野 孝伸², 齊藤 慎司³, 三好 由純¹, 松本 洋介⁴, 梅田 隆行¹, 宮下 幸長¹, 海老原 祐輔⁵
SEKI, Kanako^{1*}, AMANO, Takanobu², SAITO, Shinji³, MIYOSHI, Yoshizumi¹, MATSUMOTO, Yosuke⁴, UMEDA, Takayuki¹, MIYASHITA, Yukinaga¹, EBIHARA, Yusuke⁵

¹ 名古屋大学太陽地球環境研究所, ² 名古屋大学理学研究科, ³ 情報通信研究機構, ⁴ 千葉大学大学院理学研究科, ⁵ 京都大学生存圏研究所

¹STEL, Nagoya University, ²Dept. Physics, Nagoya University, ³NICT, ⁴Grad. Sch. of Science, Chiba University, ⁵RISH, Kyoto University

Electron acceleration mechanisms to cause drastic variation of the Earth's outer radiation belt is one of key issues of the geospace researches. While the radial diffusion of the electrons driven by ULF waves has been considered as one of the candidate mechanisms, efficiency of the mechanism under realistic ULF characteristics and distribution is far from understood. GEMSIS (Geospace Environment Modeling System for Integrated Studies) of STEL, Nagoya University, is the observation-based modeling project for understanding energy and mass transportation from the Sun to the Earth in the geospace environment. Aiming at understanding the dynamics of the inner magnetosphere during the geospace storms, the GEMSIS-Magnetosphere working team has developed a new physics-based model for the global dynamics of the ring current (GEMSIS-RC model). The GEMSIS-RC model is a self-consistent and kinetic numerical simulation code solving the five-dimensional collisionless drift-kinetic equation for the ring-current ions in the inner-magnetosphere coupled with Maxwell equations.

We applied the GEMSIS-RC model for simulation of global distribution of ULF waves to test its capability of describing fast time scale phenomena like SCs and ULF waves. Two cases of background profile, i.e., cases without/with plasmopause in the simulation domain, are compared. The result shows that existence of plasmopause strengthens ULFs outside the plasmopause and widens the MLT region where the E_r (toroidal) component is excited from initially-given E_ϕ (poloidal) component. Comparison between runs with/without ring current ions show that the existence of hot ring current ions can deform and amplify the original sinusoidal waveforms. The deformation causes the energy cascade to higher frequency range (Pc4 and Pc3 ranges). The cascade is more pronounced in the high beta case. Combination with GEMSIS-RB model reproduced rapid radial transport by the drift resonance for ions with drift period of 600 seconds as theoretically expected.

キーワード: 内部磁気圏, リングカレント, 放射線帯, ULF 波動, Pc5, ドリフト共鳴

Keywords: inner magnetosphere, ring current, radiation belt, ULF wave, Pc5, drift resonance

磁気嵐に伴う環状電場の発生とプラズマおよび高エネルギー粒子の挙動 Behaviors of Plasma and High Energy Electron Distributions During Magnetic Storms

大家 寛^{1*}

OYA, Hiroshi^{1*}

¹ 東北大学 理学研究科

¹Department of Geophysics Tohoku University

1989年2月に準極軌道に入った“あけぼの”衛星においてPWS観測器に基づく波動スペクトルに示されるUHR周波数より、プラズマ圏プラズマ密度分布は磁気嵐にともなう激しい変動を示すことが明らかになっていた。このダイナミックな変動は地磁気変動に呼応して生ずる誘導電場によるドリフトとして理解された。即ち磁気嵐主相ではRing Currentの増大と同期して発生する東向き環状電場が生まれ、この環状電場によってプラズマは動径方向外側にドリフトを開始する。他方回復相に入るとRing Currentの減少と同期して起こる西向き環状電場によってプラズマは動径方向内側に向かってドリフトをする。この際高温希薄な外部より侵入するプラズマはプラズマ圏における低温高密度プラズマと不連続面を形成して接触する様相が観測されている。

本研究ではGOSATにより観測された高エネルギー電子の総監データに対しこの環状電界の存在を検証した。その結果環状電界の存在によって帰結する挙動を提言するものである。即ちプラズマ圏と重なる放射線外帯での高エネルギー電子の挙動と分布状態は、磁気嵐にともなう環状電場の生成にともない、Ring Currentに貢献しない400keV以上の程度のフラックスを示す成分は、磁気嵐主相において動径方向外側にドリフトを開始し、逆に回復相に入るとRing Currentの減少と同期して起こる西向き環状電場によって動径方向内側に向かってドリフトする。なおこの挙動は磁場変動率にも依存し、磁気嵐毎にドリフトのエネルギーレンジの限界値に変化を示す。平均変化率80nT/20hourの場合600keV付近にみられる限界は平均変化率170nT/3hourの場合300keV前後まで下がってくる。

キーワード: プラズマ圏プラズマ, 誘導電場, 高エネルギー電子, 環状電場, 磁気嵐主相, 磁気嵐回復相

Keywords: Plasmaspheric Plasma, Induction Electric Field, High Energy Electrons, Ring Feature Electric Field, Main Phase of Magnetic Storm, Recovery Phase of Magnetic Storm

Rapid radiation belt losses occurring during high speed solar wind stream driven storms: loss mechanisms Rapid radiation belt losses occurring during high speed solar wind stream driven storms: loss mechanisms

Craig Rodger^{1*}, Aaron T. Hendry¹, Steven K. Morley², Mark A. Clilverd³, Neil R. Thomson¹, Tero Raita⁴
RODGER, Craig^{1*}, Aaron T. Hendry¹, Steven K. Morley², Mark A. Clilverd³, Neil R. Thomson¹, Tero Raita⁴

¹University of Otago, Dunedin, New Zealand, ²Los Alamos National Laboratory, USA, ³British Antarctic Survey, Cambridge, UK, ⁴SGO, University of Oulu, Finland

¹University of Otago, Dunedin, New Zealand, ²Los Alamos National Laboratory, USA, ³British Antarctic Survey, Cambridge, UK, ⁴SGO, University of Oulu, Finland

Despite being discovered at the dawn of the space age, there are still fundamental questions concerning the acceleration and loss of highly energetic electrons; Energetic electron fluxes can increase or decrease by several orders of magnitude on time scales of less than a day. The coupling of the Van Allen radiation belts to the Earth's atmosphere through precipitating particles is an area of intense scientific interest, principally due to two differing research activities: 1) radiation belt physics and 2) the response of the atmosphere to precipitating particles.

Recently, more evidence has been found for the highly dynamic nature of the outer radiation belt electron fluxes. Studies undertaken by Steve Morley and coauthors combined observations from as many as 9 particle detectors flying onboard GPS spacecraft to show a rapid losses of energetic (>230 keV) electrons from the outer radiation belt ($4 < L^* < 6$). These electron flux "dropout" were associated with a geomagnetic storm triggered by the arrival of a high speed solar wind stream interface which separates the leading slow solar wind and trailing fast solar wind. A superposed epoch analysis was undertaken of this data around the arrival of 67 solar wind stream interfaces (SWSIs), showing a strong repeatable "signal" of a rapid electron flux dropout. While the SWSI triggered geomagnetic storms, these are comparatively small ($Dst = -40$ nT, $Kp=4$) showing that even small geomagnetic disturbances can lead to major radiation belt changes. Candidate causes for the dropouts which have been put forward are losses through the magnetopause (by either magnetopause shadowing or outward diffusion) and energetic electron precipitation (EEP) into the atmosphere due to wave-particle interactions.

In this talk we will test the potential dropout causes by examining observations from multiple low-Earth orbit spacecraft, as well as ground-based instruments which are part of the AARDDVARK network.

キーワード: radiation belt, geomagnetic storms, solar wind interactions, energetic electron precipitation

Keywords: radiation belt, geomagnetic storms, solar wind interactions, energetic electron precipitation

Recent observations of dropouts of the electron radiation belt and their implications Recent observations of dropouts of the electron radiation belt and their implications

Steven Morley^{1*}

MORLEY, Steven^{1*}

¹Los Alamos National Laboratory

¹Los Alamos National Laboratory

Predicting losses or enhancements of relativistic electron fluxes during storm time has proven difficult, with only ~50% of Dst storms showing a net increase in fluxes and about 20% showing a net decrease. Under high-speed stream driving the predictability of net flux increases is enhanced by the Russell-McPherron effect but substantial variability is still observed. High-speed solar wind streams, which have their origins in the outflow from coronal holes, form a large-scale repeatable driver for Earth's magnetosphere and the parts of geospace coupled to it. Up to 13 on-orbit dosimeters in the GPS constellation can be used to monitor the global state and evolution of the electron radiation belt at unprecedented temporal and spatial resolution. Recent observations using the GPS constellation have shown that dropouts extending to $L \sim 4$ are a consistent response to high-speed streams, irrespective of whether the stream drives a geomagnetic storm. Case studies have shown that these dropouts can occur on time scales of less than 3 hrs and recovery to pre-event count levels can take in excess of two weeks. Combining these observations with the modeling framework provided by the Dynamic Radiation Environment Assimilation Model (DREAM) gives additional capabilities to understand the variability of the radiation belt and the dominant processes. In particular, we focus on results of a statistical study of a set of prolonged electron dropouts observed with GPS, combined with insight from detailed case studies and modeling. What we have learned, and how we can extend this work, will be discussed.

キーワード: radiation belt, magnetosphere, space weather

Keywords: radiation belt, magnetosphere, space weather

Key Issues in Substorm Onset and Expansion Key Issues in Substorm Onset and Expansion

Chio Z. Cheng^{1*}

CHENG, Chio Z.^{1*}

¹Plasma and Space Science Center, National Cheng Kung University

¹Plasma and Space Science Center, National Cheng Kung University

Some key issues for understanding substorms are (1) substorm onset mechanism that includes the structure of growth phase magnetosphere, the structure and breakup of auroral onset arcs; and the structure of Pi2 waves and associated instabilities; (2) substorm expansion that includes depolarization and cross-tail current disruption and associated energetic particle injections in the ring current region. Observationally there has been much progress in the last few years on the pressure profile and magnetic field structure in the cross-tail current region, the structure of bright spots along the breakup arc, and the structure of Pi2 waves in the cross-tail current region and under the onset arcs, magnetic field dipolarization, and energetic particle injection. In this talk we present key features of substorm onset observation and their theoretical explanation. We will also discuss the possible mechanism of current disruption, magnetic field dipolarization and energetic particle injection during the substorm expansion phase.

キーワード: substorm, ring current, energetic particle injection, dipolarization, current disruption

Keywords: substorm, ring current, energetic particle injection, dipolarization, current disruption

Oxygen ion flux variation associated with a substorm on May 6, 1988 Oxygen ion flux variation associated with a substorm on May 6, 1988

Kazue Takahashi^{1*}, Masahito Nose², Kunihiro Keika³
TAKAHASHI, Kazue^{1*}, NOSE, Masahito², Kunihiro Keika³

¹Johns Hopkins University Applied Physics Laboratory, ²Kyoto University, ³New Jersey Institute of Technology

¹Johns Hopkins University Applied Physics Laboratory, ²Kyoto University, ³New Jersey Institute of Technology

The flux of O⁺ ions is known to increase in the inner magnetosphere during geomagnetically active periods, but the mechanism for the flux change is not well understood. To gain insight into the mechanism we examine the variation of the O⁺ flux observed with the Charge-Energy-Mass Spectrometer (CHEM) on the AMPTE/CCE spacecraft in association with a substorm on May 6, 1988. The substorm onset occurred at ~0607 UT during the main phase of a geomagnetic storm that reached the Dst minimum of -160 nT. CCE detected a clear dipolarization ($\Delta B \sim 140$ nT over $\Delta T \sim 60$ s) at $L \sim 4.0$, magnetic latitude ~ 6 degrees, and magnetic local time ~ 20 hr. After the dipolarization, the power spectral density of broadband magnetic field variation in the Pc1 band was ~ 1.5 orders of magnitude higher than the pre-onset level. In addition, there was a brief period of narrowband pulsation at a frequency about one half of the local proton cyclotron frequency. Across the dipolarization the O⁺ flux increased by 0.5-1.0 orders of magnitude at energies below 100 keV. No flux enhancement was observed for O⁺ at energies above 100 keV and there was no obvious flux change for H⁺ over the entire energy band, 1-300 keV, covered by the CHEM instrument. We add observations by the nearby satellites GOES-6, GOES-7, and DE-1 to discuss possible mechanism(s) for the O⁺ enhancement.

キーワード: substorm, oxygen ions, inner magnetosphere

Keywords: substorm, oxygen ions, inner magnetosphere

放射線帯外帯に対するサブストームのインパクト Direct impact of substorm on outer radiation belt

海老原 祐輔^{1*}, 田中 高史²

EBIHARA, Yusuke^{1*}, TANAKA, Takashi²

¹ 京都大学生存圏研究所, ² 九州大学

¹RISH, Kyoto University, ²Kyushu University

We use a four-dimensional drift kinetic simulation coupled with a global magnetohydrodynamics (MHD) simulation to investigate the direct impact of a substorm on relativistic electrons in the heart of the radiation belt. During the substorm growth phase, energetic electrons moved earthward from the plasma sheet because the dawn-dusk convection electric field overcame the dusk-dawn induction electric field. During the expansion phase, energetic electrons traversed the quasi-trapping region and encountered the stable trapping region due to persistent and variable electric field, which is dominated by the westward component on the nightside. Finally, relativistic electrons initially located at $L \sim 4$ were replaced by newly injected ones within 1 hour. We will discuss the generation mechanism of the persistent and variable electric field that appear during the expansion phase and formation of the outer radiation belt.

キーワード: 放射線帯, サブストーム, 電子, 内部磁気圏

Keywords: Radiation belt, Substorm, energetic electrons, inner magnetosphere

ホイッスラーモード・コーラス放射との相互作用による相対論的電子フラックスの形成過程

Formation process of relativistic electron flux through interaction with whistler-mode chorus emissions

大村 善治^{1*}, 宮下 優¹, 吉川 真登¹, Summers Danny², 疋島 充³, 海老原 祐輔¹

OMURA, Yoshiharu^{1*}, MIYASHITA, Yu¹, YOSHIKAWA, Masato¹, SUMMERS, Danny², HIKISHIMA, Mitsuru³, EBIHARA, Yusuke¹

¹ 京都大学生存圏研究所, ² ニューファンドランド・メモリアル大学, ³ 東北大学

¹RISH, Kyoto University, ²Memorial University of Newfoundland, ³Tohoku University

We perform test particle simulations of energetic electrons interacting with whistler-mode chorus emissions. We first obtain a spatial and temporal profile of a chorus wave packet based on the nonlinear wave growth theory for parallel propagation. We assume the lower band chorus emissions by terminating the rising tone at half the electron cyclotron frequency because of the nonlinear damping due to the quasi-oblique propagation. We then calculate trajectories of a large number of electrons with the same initial energy and pitch angle launched at different locations of the magnetic field line and different timings with respect to a pair of chorus emissions generated at the magnetic equator. By casting the pitch angles after the interaction to the equator, and by taking into account the phase space volume including the bounce motion along the magnetic field line, we compute the equatorial distribution function of the electrons started from the delta function. The distribution function can be regarded as a numerical Green's function for one cycle of chorus wave-particle interaction. We obtain the Green's function for the energy range 10 keV - 6 MeV and all pitch angles greater than the loss cone angle. By taking the convolution integral of the Green's functions and the distribution of the injected electrons, we can obtain the time evolution of the distribution function due to interaction with one cycle of chorus emissions. Repeating the convolution integral of the Green's functions and the evolving distribution function, we can follow a long time evolution of the electron distribution function in the radiation belts. We find that the effective acceleration processes of the relativistic turning acceleration and the ultra-relativistic acceleration through interaction with chorus emissions contribute to a rapid formation of the relativistic MeV electron flux within an hour after the injection of tens of keV electrons.

Keywords: whistler-mode chorus, simulation, magnetosphere, nonlinear, relativistic electrons, radiation belts

高速太陽風通過時のホイッスラー波動励起のIMF依存性に関するシミュレーション Simulation on the IMF Bz control of the whistler mode wave excitation associated with the high-speed

三好 由純^{1*}, JORDANOVA Vania², 片岡 龍峰³, 松本 洋介⁴, 加藤 雄人⁵

MIYOSHI, Yoshizumi^{1*}, Vania Jordanova², KATAOKA, Ryuhō³, MATSUMOTO, Yosuke⁴, KATOH, Yuto⁵

¹ 名古屋大学太陽地球環境研究所, ² ロスアラモス国立研究所, ³ 東京工業大学, ⁴ 千葉大学大学院理学研究科, ⁵ 東北大学大学院理学研究科地球物理学専攻

¹STEL, Nagoya University, ²Los Alamos National Laboratory, ³Tokyo Institute of Technology, ⁴Chiba University, ⁵Tohoku University

The outer belt electron flux tends to increase largely during the high-speed coronal hole streams. The flux enhancements depend on not only the solar wind speed but also the IMF Bz; the southward-dominant streams (SBz-HSS) cause the large flux enhancement rather than the northward-dominant streams (NBz-HSS) [Miyoshi and Kataoka, 2008]. Considering the acceleration process via whistler-wave particle interactions, key parameters including hot electrons, plasmasphere, and whistler mode waves should have similar differences between SBz/NBz-HSS, and we have confirmed these differences by the statistical analysis. In this study, we will simulate the relativistic-RAM electron model [Jordanova and Miyoshi, 2005] to confirm these parameter dependences on the key parameters. Firstly, we prepared the superposed LANL/MPA data about SBz/NBz-HSS as a boundary condition at L=6.6 to model the differences of the plasma sheet between SBz/NBz-HSS. In the SBz-HSS, we observe enhancements of hot electrons of ~30 keV and lower-band whistler mode waves around L=4 at dawn-side. On the other hand, in the NBz-HSS, the enhancements of ~30 keV electrons and whistler mode waves are observed at L>5. These differences primarily come from the magnetospheric convection. The betatron acceleration of the convective transport is the main driver to generate the whistler mode waves. We found that the plasma sheet temperature in the SBz-HSS is higher than that in the NBz-HSS, but the differences of the plasma sheet temperature do not cause clear differences of the whistler mode wave excitation between SBz/NBz-HSS.

キーワード: 内部磁気圏, シミュレーション, 太陽風磁気圏結合

Keywords: Inner Magnetosphere, simulation, solar wind - inner magnetosphere

ホイッスラーコーラスによる相対論的電子マイクロバーストについて:GEMSIS-RBWシミュレーション

Relativistic electron microbursts driven by whistler chorus: GEMSIS-RBW simulations

齊藤 慎司^{1*}, 三好 由純², 関 華奈子²

SAITO, Shinji^{1*}, MIYOSHI, Yoshizumi², SEKI, Kanako²

¹ 情報通信研究機構, ² 名古屋大学太陽地球環境研究所

¹SWEI Laboratory, NICT, ²STE Laboratory, Nagoya University

Electron microbursts, which are bursty enhancements of precipitated electron flux during several tens ms, are likely to make an important contribution to high-energy electron flux loss in the outer radiation belt. The SAMPEX satellite frequently observed that the relativistic electron microbursts ($>1\text{MeV}$) were accompanied with the lower band whistler chorus (Lorentzen et al. (2001)). The observations suggest that whistler chorus not only accelerate radiation belt electrons, but also decrease the electron flux. As discussed in Horne et al. (2003), relativistic electrons with small pitch angle close to a loss cone resonate with the lower band whistler chorus at relatively high magnetic latitudes (> 30 deg.). Considering that the whistler chorus propagates at the high magnetic latitudes, we investigate wave-particle interaction process between the whistler chorus and the relativistic electrons bouncing along the magnetic field lines using the GEMSIS-RBW code. This code is a three-dimensional relativistic test particle simulation code demonstrating electron scattering by whistler chorus in a realistic time and spatial scales. We show that a rising tone of whistler chorus scatters electrons at the high magnetic latitudes and produces a bursty enhancement of relativistic electron precipitation flux. The duration of the burst corresponds to that of the rising tone, which are several tens ms. Our simulation results suggest that whistler chorus propagating at high magnetic latitudes precipitates the relativistic electrons showing the relativistic electron microbursts.

キーワード: 波動粒子相互作用, 相対論的電子マイクロバースト, テスト粒子シミュレーション

Keywords: wave-particle interaction, relativistic electron microburst, test-particle simulation

Recurrent and propagating SAPS/ SAID structures observed by the SuperDARN Hokkaido radar

Recurrent and propagating SAPS/ SAID structures observed by the SuperDARN Hokkaido radar

西谷 望^{1*}, 小川 忠彦², Sasha Koustov³, 菊池 崇¹, 海老原 祐輔⁴, 片岡 龍峰⁵, 堀 智昭¹, 細川 敬祐⁶, 寺本 万里子¹
NISHITANI, Nozomu^{1*}, OGAWA, Tadahiko², Sasha Koustov³, KIKUCHI, Takashi¹, EBIHARA, Yusuke⁴, KATAOKA, Ryuhō⁵,
HORI, Tomoaki¹, HOSOKAWA, Keisuke⁶, TERAMOTO, Mariko¹

¹Solar-Terrestrial Environment Laboratory, Nagoya University, ²National Institute of Information and Communications Technology, ³University of Saskatchewan, ⁴Research Institute of Sustainable Humanosphere, Kyoto University, ⁵Tokyo Institute of Technology, ⁶University of Electro-Communications

¹Solar-Terrestrial Environment Laboratory, Nagoya University, ²National Institute of Information and Communications Technology, ³University of Saskatchewan, ⁴Research Institute of Sustainable Humanosphere, Kyoto University, ⁵Tokyo Institute of Technology, ⁶University of Electro-Communications

We study two events of the recurrent propagating SAPS/SAID structures at magnetic latitudes of $\sim 60^\circ$, which were observed by the Hokkaido radar at 0830 - 0930 UT (1730 - 1830 MLT) on April 23, 2008 and at 1100-1500 UT on March 1, 2011. Both events occurred during the main phases of moderate geomagnetic storms with minimum Dst about -37 nT and -75 nT respectively. The IMF was southward and turned to northward just prior to the first event, whereas it was constantly southward for the second event. For both events the period, velocity of westward progression and wavelength are 30 min, 700 m/s and 1200 km respectively. They have similar characteristics to those of the 'auroral giant undulation' at the equatorward edge of the auroral oval although some numbers, for example the periods, are slightly different. The radar data presented strongly support an idea that the recurrent SAPS / SAID structures are generated through internal processes in the inner magnetosphere, due to probably some kind of plasma instability, rather than due to externally-driven processes.

キーワード: 北海道-陸別 HF レーダー, SuperDARN, SAPS / SAID, 内部磁気圏プラズマ不安定

Keywords: Hokkaido HF radar, SuperDARN, SAPS / SAID, plasma instability in the inner magnetosphere

ヘリウムブランチ電磁イオンサイクロトロントリガード放射による高エネルギー粒子降下

Precipitation of Highly Energetic Protons by Helium Branch Electromagnetic Ion Cyclotron Triggered Emissions

小路 真史^{1*}, 大村 善治¹

SHOJI, Masafumi^{1*}, OMURA, Yoshiharu¹

¹ 京都大学生存圏研究所

¹ Research Institute for Sustainable Humanosphere, Kyoto University

In the Earth's inner magnetosphere, the electromagnetic ion cyclotron (EMIC) triggered emissions scatter energetic protons around the equatorial region. We perform a one-dimensional hybrid simulations with cylindrical parabolic magnetic geometries to investigate the precipitation of energetic protons controlled by the EMIC triggered emissions in the He⁺ branch. We generate the EMIC triggered emissions in the He⁺ branch of the dispersion relations around the equatorial region of the simulation space. The nonlinear wave growth theory shows a good agreement with the simulation result. Due to the difference of the resonance velocity between the H⁺ and He⁺ branch triggered emissions, the He⁺ branch triggered emissions scatter highly energized protons around the equatorial region, resulting in the precipitations of the particles into the polar regions. Estimation of the kinetic energy of the resonant particles predicts that the energy reaches around 3 MeV outside the plasmopause although the H⁺ branch triggered emissions can interact with the particles with 1MeV.

キーワード: EMIC トリガード放射, ハイブリッドシミュレーション, プロトン降下, 非線形波動粒子相互作用, 内部磁気圏
Keywords: EMIC triggered emission, hybrid simulation, proton precipitation, nonlinear wave particle interaction, inner magnetosphere

Geotail 衛星観測に基づくコーラスエミッションの非線形成長特性の解析 Evaluation of nonlinear growth of chorus emissions observed by Geotail

森 晋作¹, 幅岸 俊宏¹, 八木谷 聡^{1*}, 大村 善治², 小嶋 浩嗣²

MORI, Shinsaku¹, HABAGISHI, Toshihiro¹, YAGITANI, Satoshi^{1*}, OMURA, Yoshiharu², KOJIMA, Hirotsugu²

¹ 金沢大学, ² 京都大学生存圏研究所

¹Kanazawa University, ²RISH, Kyoto University

Geotail 衛星観測による電磁界波形データを用いて、コーラスエミッションの発生及び伝搬特性解析を行う。コーラスエミッションの非線形成長理論において、振幅の変化と相関を持つ周波数変化が報告されている [1]。その特徴はシミュレーションで再現されたコーラスエミッションでも確認され、理論の妥当性が示唆された [2][3]。今回の解析では、Geotail 衛星により実際に地球磁気圏で観測されたコーラスエミッションにおいて、非線形成長の特徴がみられるかどうかを検証すると共に、伝搬過程における振幅の成長及び 1/2 サイクロトロン周波数での減衰についても考察した。

地球磁気圏の昼側 (L 値が 9~10 の領域) で観測されたコーラスエミッションの磁界波形データを用いて解析を行った。非線形成長理論では“ 振幅 × 周波数 ”と“ 周波数変化率 ”の関係性が示されているため [1]、Geotail 衛星が観測したコーラスエミッションに対し、それらの相関関係を検証した。波形データから振幅と周波数の時間変化を算出するため、解析信号法 (ASM: Analytic Signal Method) を適用し、コーラスエミッションの周波数変化に通過帯域が追従するバンドパスフィルタを用い、背景ノイズを除去して解析を行った。相関解析の結果、非線形成長理論で予測される通り、“ 振幅 × 周波数 ”と“ 周波数変化率 ”の間に正の相関関係が見られた。

また、1/2 サイクロトロン周波数のギャップが明確に生じているコーラスエミッションに対して、その発生領域、及び伝搬過程における振幅の変化について解析した。ローワーバンドコーラスにおける周波数スペクトルの上端はコーラスエミッションの発生領域における 1/2 サイクロトロン周波数、またアッパーバンドコーラスにおける周波数スペクトルの下端はコーラスエミッションの観測領域における 1/2 サイクロトロン周波数を表すと考えられる。実際、Tsyganenko の地球磁場モデルを用いると、ローワーバンドコーラスにおける周波数スペクトルの上端が、衛星観測点につながる磁力線に沿って地磁気強度が最小となる領域の 1/2 サイクロトロン周波数と等しくなり、ここがコーラスの発生領域である可能性が示された。一方、発生領域におけるコーラスエミッションの振幅を非線形成長理論から推定し、発生領域から観測点までの伝搬における非線形成長を評価した [2][3]。その結果、観測されたコーラスの振幅は、非線形成長理論から予測される値と矛盾ないことが分かった。一方で、電子がコーラスエミッションの電界成分に捕捉されることによる 1/2 サイクロトロン周波数での減衰 [4] についても、観測データによる検証を行った。

[1] Omura Y., Y. Katoh, and D. Summers (2008), Theory and simulation of the generation of whistler-mode chorus, *J. Geophys. Res.*, 113, A04223, doi:10.1029/2007JA012622.

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[4] Omura, Y., M. Hikishima, Y. Katoh, D. Summers, and S. Yagitani (2009), Nonlinear mechanisms of lower-band and upper-band VLF chorus emissions in the magnetosphere, *J. Geophys. Res.*, 114, A07217, Doi:10.1029/2009JA014206.

キーワード: Geotail 衛星, コーラスエミッション, 非線形成長, 振幅, 周波数変化率, 発生 伝搬

Keywords: Geotail spacecraft, Chorus emission, Nonlinear growth, Amplitude, Frequency sweep rate, Generation and propagation

カナダ・アサバスカでのLF帯電波観測による高エネルギー電子降下現象観測：初期観測結果

LF radio observation of storm-time energetic electron precipitation at Athabasca: Initial result

土屋 史紀^{1*}, 三澤 浩昭¹, 森岡 昭¹, 小原 隆博¹, 三好 由純², 塩川 和夫², 菊池 崇², マーチンコナーズ³

TSUCHIYA, Fuminori^{1*}, MISAWA, Hiroaki¹, MORIOKA, Akira¹, OBARA, Takahiro¹, MIYOSHI, Yoshizumi², SHIOKAWA, Kazuo², KIKUCHI, Takashi², Martin Connors³

¹ 東北大学大学院理学研究科, ² 名古屋大学太陽地球環境研究所, ³ アサバスカ大学地球物理学観測所

¹Graduate School of Science, Tohoku University, ²Solar-Terrestrial Environment Laboratory, Nagoya University, ³Athabasca University Geophysical Observatory

Low frequency (LF) radio observation system has been installed in Athabasca, Canada at Oct. 2010. Purpose of the observation is to investigate energetic electron precipitation from outer radiation belts to the atmosphere in the sub-auroral latitude region during magnetic storms and substorms. In this paper, some initial results will be presented. LF transmitter signal perturbation is caused by ionization of lower edge of the ionosphere and sensitive to higher energy electron (>100keV) precipitation than riometer observations. In 2011, the electron precipitation signatures were detected during the main and early recovery phases in several magnetic storms. During the night time, fluctuations in the LF phase data with the time scale of Pc5 or longer were found. Comparison of the phase fluctuation with the GOES magnetic field data during a small magnetic storm on 5 June 2011 shows significant correlation, which implies the drift resonance of energetic electrons and/or Pc5 modulation of energetic electron scattering efficiency in the inner magnetosphere.

On the evolution of nonlinear ion acoustic waves and their stability in an unmagnetized plasma

On the evolution of nonlinear ion acoustic waves and their stability in an unmagnetized plasma

Amar Kakad^{1*}, Yoshiharu Omura¹

KAKAD, Amar^{1*}, OMURA, Yoshiharu¹

¹RISH, Kyoto University, Kyoto, JAPAN.

¹RISH, Kyoto University, Kyoto, JAPAN.

One dimensional fluid simulation has been performed to study the evolution of nonlinear ion acoustic waves and their stability in an unmagnetized plasma. A Gaussian perturbation is used to model the initial localized density perturbation. It is found that the time evolution of such a localized density perturbation evolve into ion acoustic solitons, as predicted by the fluid theory. Issues pertaining to their stability and mutual collisional interaction will be addressed. Possible applications to the electrostatic solitary structures observed in some of the regions of Earth magnetosphere will be discussed.

キーワード: Ion acoustic waves, Electrostatic solitary structures, Fluid Simulation, Ion acoustic solitons, Magnetosphere

Keywords: Ion acoustic waves, Electrostatic solitary structures, Fluid Simulation, Ion acoustic solitons, Magnetosphere

SPRINT-B/ERG 衛星の超高エネルギー電子観測装置 The Extremely High-Energy Plasma/Particle Sensor for Electron of the SPRINT-B/ERG satellite

東尾 奈々^{1*}, 松本 晴久¹

HIGASHIO, Nana^{1*}, Matsumoto Haruhisa¹

¹ 宇宙航空研究開発機構 研究開発本部宇宙環境グループ

¹Japan Aerospace Exploration Agency

The ERG satellite is a science mission of JAXA/ISAS, the purpose of which is to reveal the acceleration mechanism of high energy particles in the radiation belt. The high energy electrons of radiation belts change dramatically during a magnetic storm. After such storms occur, these high energy electrons disappear, only to subsequently proliferate during the recovery phase. There are said to be two hypotheses concerning this process, external supply (adiabatic process) and internal acceleration (non-adiabatic process). The ERG satellite consists of four instrument parts, Plasma/Particle (PPE), geomagnetic field (MGF), Plasma wave (PWE) and electric field (PWE). The PPE sensor consists of five particle sensors and four plasma sensors. Our group is now developing one of the sensors, namely the eXtremely high-Energy Plasma/particle sensor for electron (XEP-e). The XEP-e observes 100keV~20MeV electrons and has four solid-state silicon detectors (SSDs) and a high-Z scintillator (GSO). It has one way conic sight and an electric part is unified with a part of sensor that is covered with aluminum to protect from contamination. The front part of the SSDs discriminate a radiation enters into the sensor and the back part of the plastic scintillator get the value of its energy. Since this satellite will traverse radiation belts, we have to assume a harsh radiation environment. Our group has an electrostatic accelerator in Tsukuba Space Center, via which we can make a radiation environment in space and proof a sensor. It is a pelletron charging system, which involves the accelerator radiating electrons or protons to a sensor in a vacuum chamber. The energies of the electrons and protons are 0.4 ~ 2.0MeV and the beam current is 1fA ~ 10nA (table 4). If the beam current of the electrons is set to 1pA, the flux becomes almost 1×10^6 counts/ sec/cm². It also has an electron-gun (5 ~ 50keV). We are using this system to proof the XEP-e and now we are still developing .

Keywords: XEP, electron, Sprint-B/ERG