Japan Geoscience Union Meeting 2013

(May 19-24 2013 at Makuhari, Chiba, Japan)

©2013. Japan Geoscience Union. All Rights Reserved.



PEM26-P02

会場:コンベンションホール

## 太陽風中で観測される上流伝搬ホイッスラーモード波動の類似性 Similarities of upstream whistler-mode waves observed in the solar wind

津川 靖基<sup>1\*</sup>, 加藤 雄人<sup>1</sup>, 寺田 直樹<sup>1</sup>, 小野 高幸<sup>1</sup>, 綱川 秀夫<sup>2</sup>, 高橋 太<sup>2</sup>, 渋谷 秀敏<sup>3</sup>, 清水 久芳<sup>4</sup>, 松島 政貴<sup>2</sup> Yasunori Tsugawa<sup>1\*</sup>, Yuto Katoh<sup>1</sup>, Naoki Terada<sup>1</sup>, Takayuki Ono<sup>1</sup>, Hideo Tsunakawa<sup>2</sup>, Futoshi Takahashi<sup>2</sup>, Hidetoshi Shibuya<sup>3</sup>, Hisayoshi Shimizu<sup>4</sup>, Masaki Matsushima<sup>2</sup>

## <sup>1</sup> 東北大学理学研究科地球物理学専攻,<sup>2</sup> 東京工業大学大学院理工学研究科地球惑星科学専攻,<sup>3</sup> 熊本大学大学院自然科学研究科,<sup>4</sup> 東京大学地震研究所

<sup>1</sup>Department of Geophysics, Tohoku University, <sup>2</sup>Department of Earth and Planetary Sciences, Tokyo Institute of Technology, <sup>3</sup>Department of Earth and Environmental Sciences, Kumamoto University, <sup>4</sup>Earthquake Research Institute, University of Tokyo

Narrowband whistler-mode waves whose frequencies close to 1 Hz have been observed near the Moon by WIND [Farrell et al., 1996], Geotail [Nakagawa et al., 2003], Lunar Prospector [Halekas et al., 2006, 2008], and Kaguya spacecraft [Tsugawa et al., 2011]. These waves are propagated upstream against the solar wind and Doppler-shifted in the spacecraft frame to be left-hand polarized. Similar waves have been observed in the upstream regions of many solar system bodies over four decades and have been called '1 Hz waves' [e.g., Heppner et al., 1967; Russell et al., 1971; Fairfield, 1974; Orlowski et al., 1990, 1995; Brain et al., 2002]. However, some unclear issues in their propagation and generation processes have not been solved, such as required condition to observe the waves, spectral formation mechanism, and the frequency dependences of the processes.

In the present study, we compare the waves in different regions to answer the issues. We investigate common properties of the waves observed by Kaguya at 100 km altitude of the Moon and by Geotail in the upstream region within 20  $R_E$  from the Earth's bow shock. Group velocity vectors of the waves in both regions are going to cancel the solar wind velocity vector. This condition is required to observe the waves and would make the narrowband spectra. We suggest that the wave frequency is determined so as to satisfy the condition. We reveal that the wave propagation angles in the upstream region of the Earth's bow shock are typically in the range of 20?40 degrees and are smaller than those near the Moon which are in the range of 20?70 degrees. This result suggests differences of the wave generation and damping processes.