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Accretion of Solid Materials onto Circum-Planetary Disks
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Regular satellites of the giant planets in our solar system are believed to be formed in circum-planetary disks around the planets during the final stage of the formation. Recent hydrodynamic simulations have revealed that gas disks around giant planets are inevitably formed in the course of gas accretion growth phase. However, in order to form satellites, solid materials are necessary in the gas disks and thus should also be accreted onto the circum-planetary disks from proto-planetary disks.

In this study, we performed orbital simulations of solid particles which is rotating in heliocentric orbits in order to investigate accretion efficiency onto circum-planetary disks under the influence of gas accretion flow. We found that the accretion efficiency of the solid particles peaks around 10m-sized particles because energy dissipation of drag with circum-planetary disk gas in this size regime is most effective. The efficiency for particles larger than 10m size becomes lower because gas drag becomes less effective for larger particles. For particles smaller than 10m, the efficiency is lower because the particles are strongly coupled with the back-ground gas flow, which prevent from accretion. We will discuss satellite formation process based on the obtained accretion efficiency of solid particles.