Japan Geoscience Union Meeting 2015

(May 24th - 28th at Makuhari, Chiba, Japan)

©2015. Japan Geoscience Union. All Rights Reserved.

SGL39-07

Room:A03



Time:May 24 10:30-10:45

## Investigation of martian surface history: NanoSIMS analyses of D/H ratios and U-Pb chronology of martian meteorites

KOIKE, Mizuho $^{1\ast}$ ; SANO, Yuji $^1$ ; TAKAHATA, Naoto $^1$ ; ISHIDA, Akizumi $^1$ ; SUGIURA, Naoji $^2$ ; ANAND, Mahesh $^3$ 

<sup>1</sup>Atmosphere and Ocean Research Institute, University of Tokyo, <sup>2</sup>Department of Earth and Planetary Science, Graduate School of Science, University of Tokyo, <sup>3</sup>Department of Physical Sciences, The Open University, UK

**Introduction:** Water is an important volatile for environments of terrestrial planets as well as their habitability. A number of recent studies have identified strong evidence for liquid water on past Mars, such as clay minerals and fluvial geomorphological features (e.g. [1][2]), whereas a comprehensive history of the martian environment remains complicated. The isotopic compositions of hydrogen (D/H) of present martian atmosphere is highly elevated (~5 times that of terrestrial water; [3]), which results from the extensive atmospheric escape. Martian meteorites are useful as they potentially provide valuable records including D/H of the past surface water and the mantle primitive water (e.g. [4][5][6][7][8]). However, due to their complicated history on Mars, it is challenging to understand their isotopic records accurately. Phosphates are helpful, for they can preserve both U-Pb chronology and D/H information. Here, we report D/H ratios and U-Pb ages of phosphates in two martian meteorites; an ancient orthopyroxenite, ALH 84001 (ALH), and a young enriched shergottite, LAR 06319 (LAR). In addition, we have also measured D/H ratio of melt-inclusion glass (MIs) in LAR.

**Analytical methods:** For ALH, 3 merrillite grains with known U-Pb ages [9] were selected for D/H analyses. For LAR, several phosphate grains and MIs were found in a thin section, using a SEM-EDS. Both U-Pb and D/H analyses were carried out using a NanoSIMS 50 installed at AORI, Univ. of Tokyo. Before SIMS analyses, the samples were baked at ~100 °C in a SIMS air-lock overnight before/after gold coating to remove adsorbed water. The analytical methods of U-Pb dating were the same as the previous study [9]. The D/H analyses were conducted on the phosphates in ALH and LAR and MIs in LAR. A Cs<sup>+</sup> primary ion beams with 200pA/1nA was used for phosphates and MIs, respectively. An electron gun was used for charge compensation. Negative secondary ions of  ${}^{1}\text{H}^{-}$ ,  ${}^{2}\text{D}^{-}$ ,  ${}^{12}\text{C}^{-}$  and  ${}^{18}\text{O}^{-}$  were collected. A natural terrestrial apatite from Morocco and NIST SRM 610 were used as standards. To avoid terrestrial H contamination, most of which are background H in the analysis chamber and hydrocarbon in the sample cracks, careful analytical protocols following a previous study [7] were conducted.

**Results & Discussion:** The  $\delta$ D values of ALH merrillite varied from -300 to 1970 ‰ (Fig). The obtained highest value is similar to those of ALH carbonates and maskelynite [4]. The U-Pb age of the same grains, 3990Ma, can be interpreted as an impact-induced reset age [9]. It is likely that their D/H ratios reflect 3990Ma surface water, incorporated during the impact and/or a later hydrous metamorphism. The high D/H mainly supports a previously proposed two-stage evolution [5]. On the other hand, phosphates in LAR yielded a total Pb/U age as 167+/-57 Ma. This is consistent with other radiometric ages within uncertainty, suggesting the U-Pb system in the phosphates has been preserved since crystallization of the host rock.  $\delta$ D values of LAR apatite, merrillite and MIs were 3340-4380 ‰, 1070-5260 ‰ and 1150-6830 ‰, respectively (Fig). The D/H may have mineral trends: MIs >merrillite >apatite. A similar trend was reported previously [7]. While apatite possibly recorded the magmatic water at the timing of crystallization, MIs might have incorporated water from another reservoir with extremely high D/H ratios.

References: [1] Bibring et al. (2006) Science 312, 400-404. [2] Ehlmann et al. (2011) Nature 479, 53-60. [3] Owen et al. (1988) Science 240, 1767-1770. [4] Sugiura and Hoshino (2000) Meteorit. Planet. Sci. 35, 373-380. [5] Greenwood et al. (2008) Geophys. Res. Lett. L05203, 1-5. [6] Usui et al. (2012) Earth Planet. Sci. Lett. 357, 119-129. [7] Hu et al. (2014) Geochim. Cosmochim. Acta 140, 321-333. [8] Usui et al. (2015) Earth Planet. Sci. Lett. 410, 140-151. [9] Koike et al. (2014) Geochem. J. 48, 423-431.

Keywords: martian meteorites, NanoSIMS, D/H ratios, U-Pb dating, phosphates, melt-inclusions

## Japan Geoscience Union Meeting 2015 (May 24th - 28th at Makuhari, Chiba, Japan)

©2015. Japan Geoscience Union. All Rights Reserved.

SGL39-07

Room:A03



Time:May 24 10:30-10:45

